



# Implications of the Worsening GCR Radiation Environment

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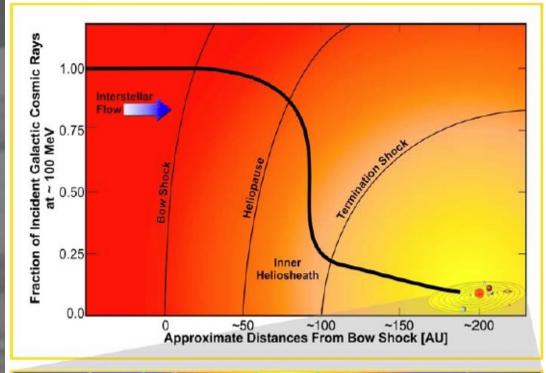


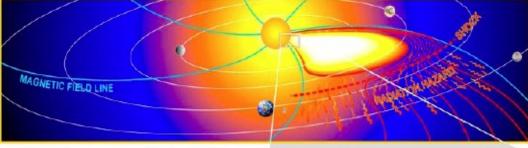


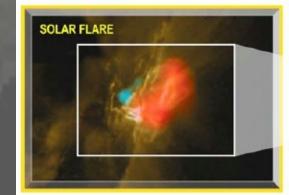


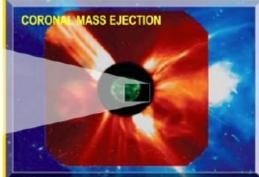
 GCRs from the Galaxy and Beyond

 SEPs from Flares and CMEs













#### **Space Weather**

#### RESEARCH ARTICLE

10.1002/2014SW001084

#### Special Section:

The Crater Special Issue of Space Weather: Building the Observational Foundation to Deduce Biological Effects of Space Radiation

#### **Key Points:**

- GCR radiation is increasingly hazardous
- Radiation limited duration for missions in deep space
- Timing during solar cycle of missions remains a critical factor

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## Does the worsening galactic cosmic radiation environment observed by CRaTER preclude future manned deep space exploration?

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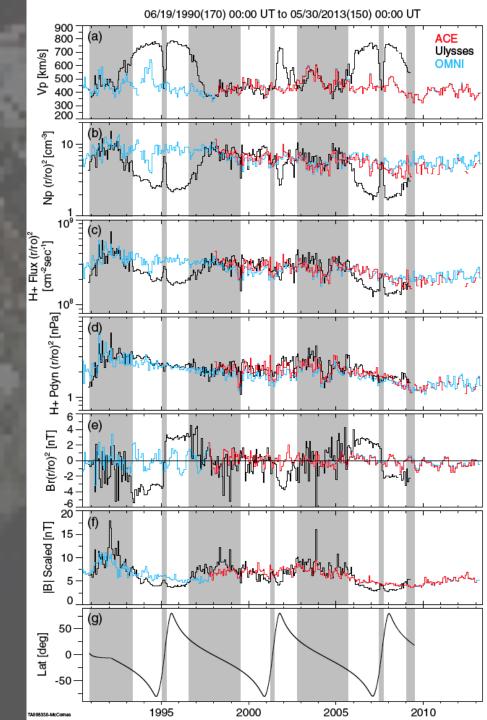
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**Abstract** The Sun and its solar wind are currently exhibiting extremely low densities and magnetic field strengths, representing states that have never been observed during the space age. The highly abnormal solar activity between cycles 23 and 24 has caused the longest solar minimum in over 80 years and continues into the unusually small solar maximum of cycle 24. As a result of the remarkably weak solar activity, we have also observed the highest fluxes of galactic cosmic rays in the space age and relatively small solar energetic particle events. We use observations from the Cosmic Ray Telescope for the Effects of Radiation (CRaTER) on the Lunar Reconnaissance Orbiter to examine the implications of these highly unusual solar conditions for human space exploration. We show that while these conditions are not a show stopper for long-duration missions (e.g., to the Moon, an asteroid, or Mars), galactic cosmic ray radiation remains a significant and worsening factor that limits mission durations. While solar energetic particle events in cycle 24 present some hazard, the accumulated doses for astronauts behind 10 g/cm<sup>2</sup> shielding are well below current dose limits. Galactic cosmic radiation presents a more significant challenge: the time

# Protracted Min (23) and Mini Max (24)

- Dropping solar wind
  - Flux
  - Pressure
  - Magnetic Field
- Continues trend observed by Ulysses

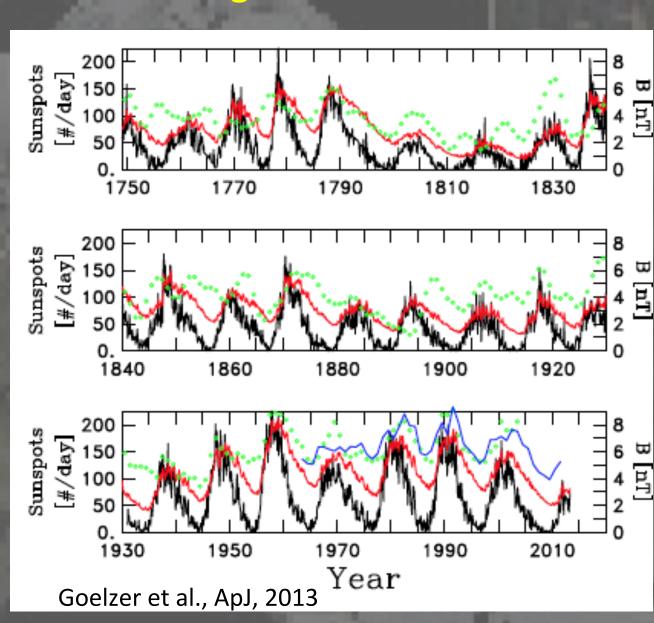
McComas et al., ApJ, 2013



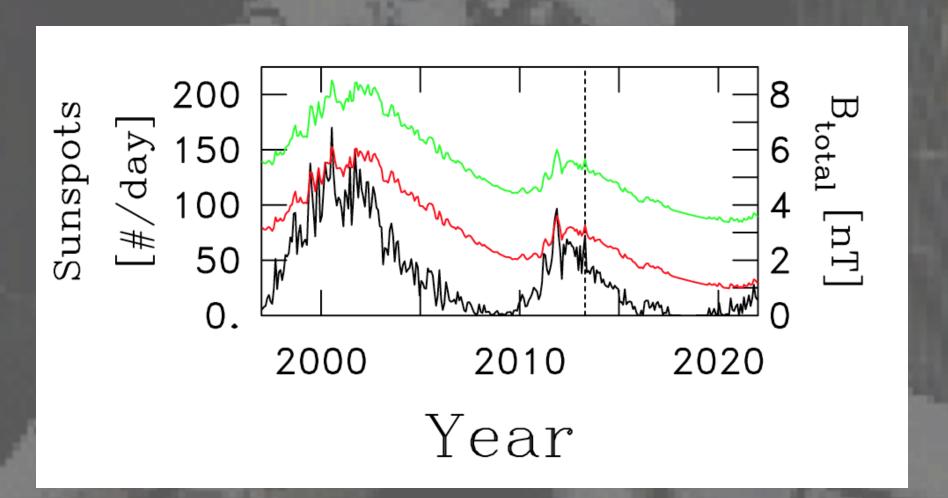
#### Long-term Record of Magnetic Field & SSN

- SSN (black)
- Predicted (red)
- OMNI (blue)
- <sup>10</sup>Be (green)
- Magnetic Flux

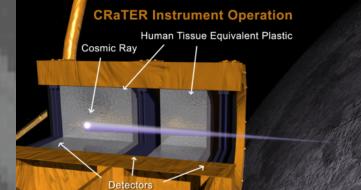
   Balance
   (Schwadron et al.,, 2010)

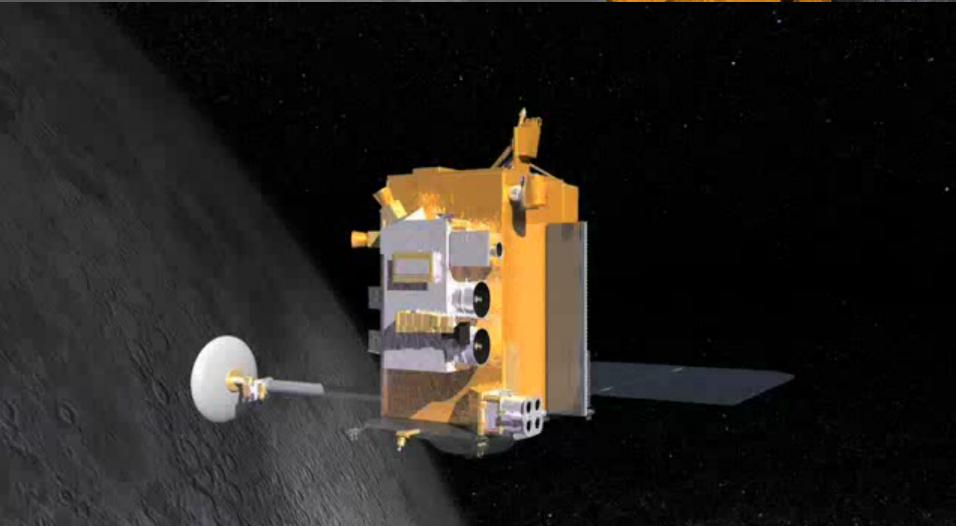


# Continued Decay of Magnetic Flux in the Dalton-like Minimum



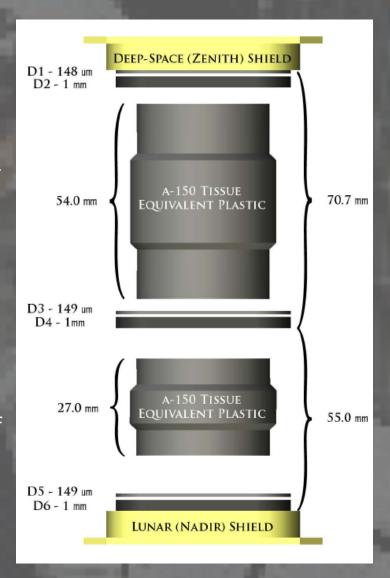
### **CRaTER Concept of Operations**





### **Quick CRaTER Overview**

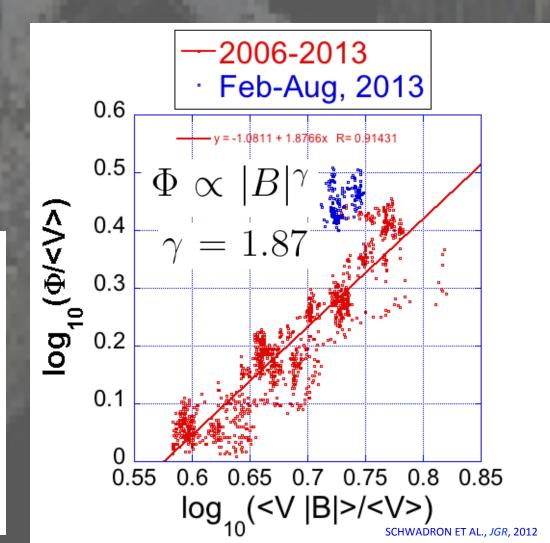
- Measures > 12 MeV/nuc particles
- 6 detectors (D1-D6) with Tissue Equivalent
   Plastic (TEP) between pairs of detectors
  - Thick and thin detectors with different gains allow a large range of Linear Energy Transfer (LET) to be sampled
  - TEP mimics absorption of energy by human tissue as radiation passed through telescope
- Senses particles from zenith and nadir directions
- Any energy deposit in any detector triggers an 'event', in which all energy deposits from all detectors are recorded
- Data products are in terms of LET, the amount of energy deposited per path-length (ΔΕ/Δx) as a particle transits through detector

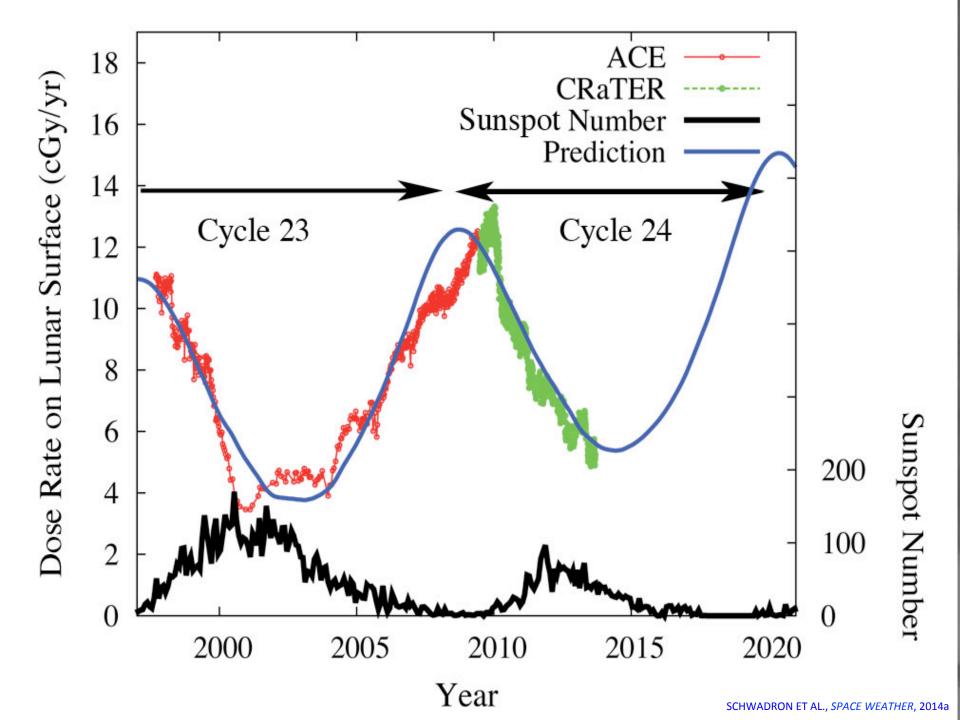


# Slab Turbulence Model for Modulation of Galactic Cosmic Rays

 Modulation Potential characterizes evolution of GCRs fluxes (protons & heavy ions)

$$\Phi = |Ze|\phi(r)$$
 
$$\phi(r) = \int_{r}^{R_{b}} dx \frac{V(x)}{3\kappa_{1}(x)}$$
 
$$\kappa_{\parallel} \propto r_{g}^{2}/F^{2} \quad F = \delta B/B$$
 
$$\Phi \propto B^{2}$$

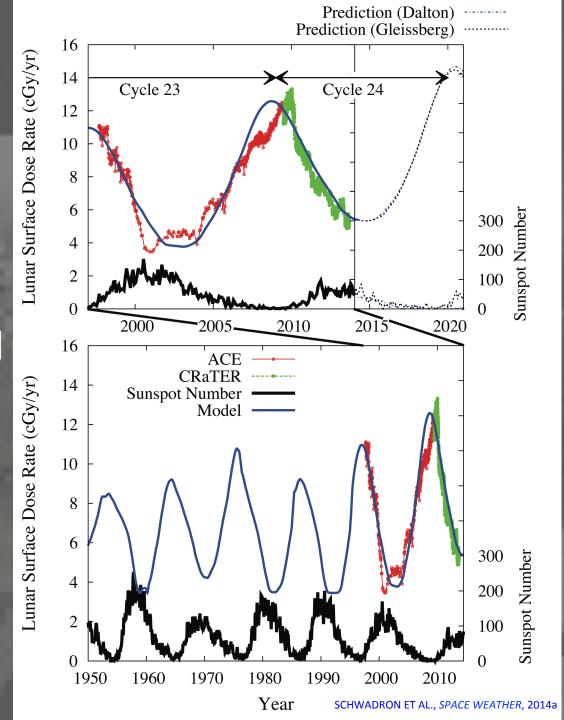




## Details about the Modeling

- GCR fluxes modeled using slab-turbulence + force-free model
  - Similar model as Badhwar O'Neill, however modulation potential scaled differently based on observations
- Utilize a three-layer version of HZETRN 2005 and HZETRN 2010 for modeling doses, doseequivalents, and effective dose

- Longer-term
   trend deduced
   from observed
   Heliospheric
   Magnetic field
- Continues
   trend observed
   by ACE and
   CRaTER



### 3% Risk for Exposure Induced Death

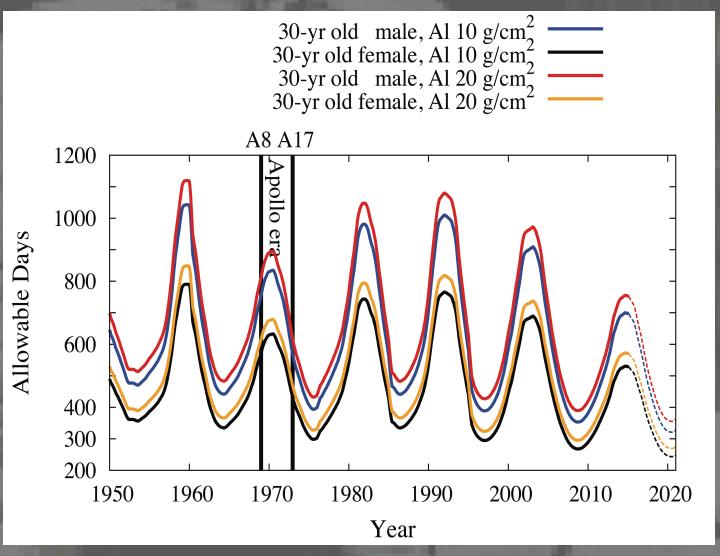
Age	3% REID Male E (cSv)	3% REID Female E (cSv)
30	62	47
45	95	75
55	147	112

Managing Space Radiation Risk in the New Era of Space Exploration Committee on the Evaluation of Radiation Shielding for Space Exploration, National Research Council 2008

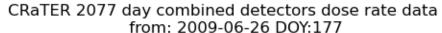
Note: E. Semones will present an updated approach to risk

# Implications of the Changing Space Weather Environment

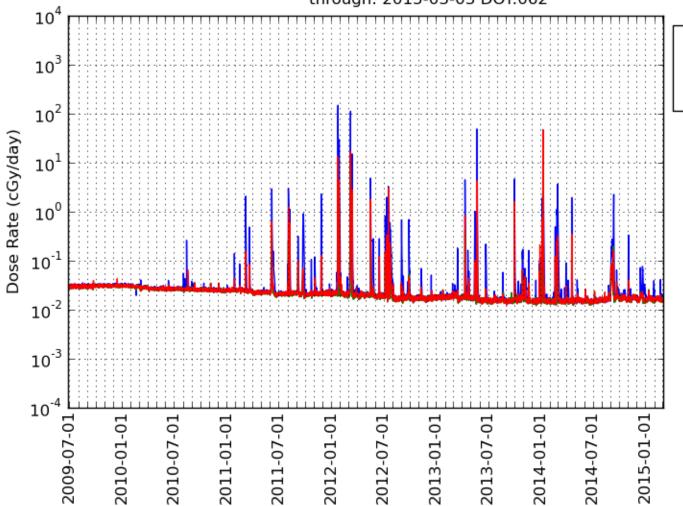
- HZETRN used to derive effective dose rate
- "Allowable days" simply the effective dose limit/ effective dose rate



## CRaTER – the complete record



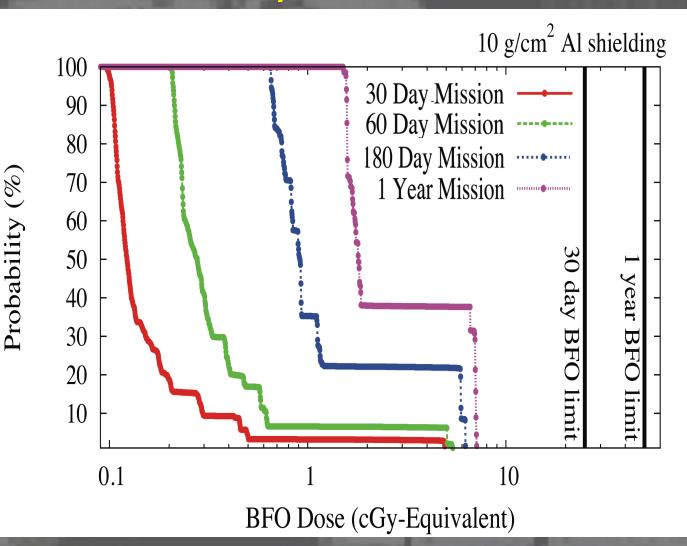
from: 2009-06-26 DOY:177 through: 2015-03-03 DOY:062



- Detectors 1 & 2
- Detectors 3 & 4
- Detectors 5 & 6

# PREDICCS Data Showing Low Probability of SEP events in cycle 23-24

- PREDICCS 
   Predictions of
   radiation from
   REleASE,
   EMMREM, and
   Data
   Incorporating
   CRaTER, COSTEP,
   and other SEP
   measurements
- On-line system to specify radiation environment through interplanetary space

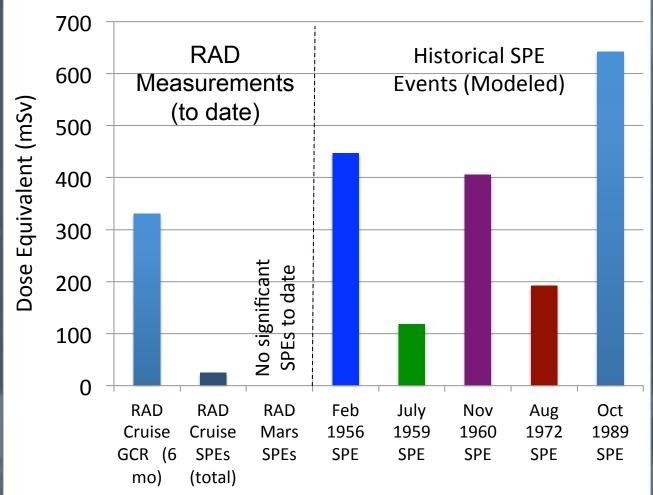




#### Comparison of Total Radiation Dose Equivalent measured by RAD to modeled Historic SPE Events



## Contribution of SPEs to Total Dose Eq. (mSv) (behind 5 g/cm<sup>2</sup> Al shielding)



\*SPE Dose Equivalent values modeled behind 5 g/cm2 Aluminum by M.-H. Kim, F. Cucinotta, et al. (AGU, 2012).

RAD cruise measurements from Jan-July 2012.

Nov. 60 SPE includes contributions from 2 events.

Oct. 89 SPE includes contributions from 5 events over 1 month.

### Approach, Assumptions

- Direct observations critical for validating models!!!
- GCR fluxes modeled using slab-turbulence + forcefree model
  - Modulation potential scaled with observed magnetic field strength
  - Dose-rate validated with ACE and CRaTER
- Three-layer version of HZETRN 2005 and HZETRN 2010 for modeling doses, dose-equivalents, and effective dose
  - LET spectra from HZETRN validated with CRaTER









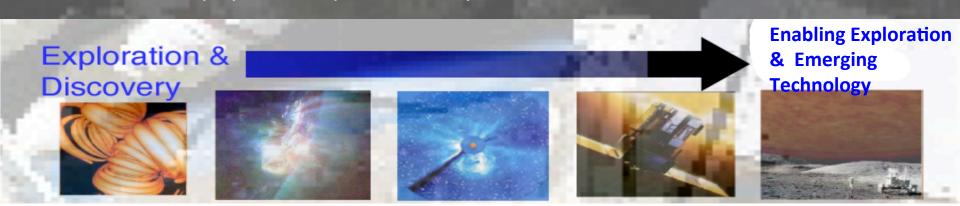


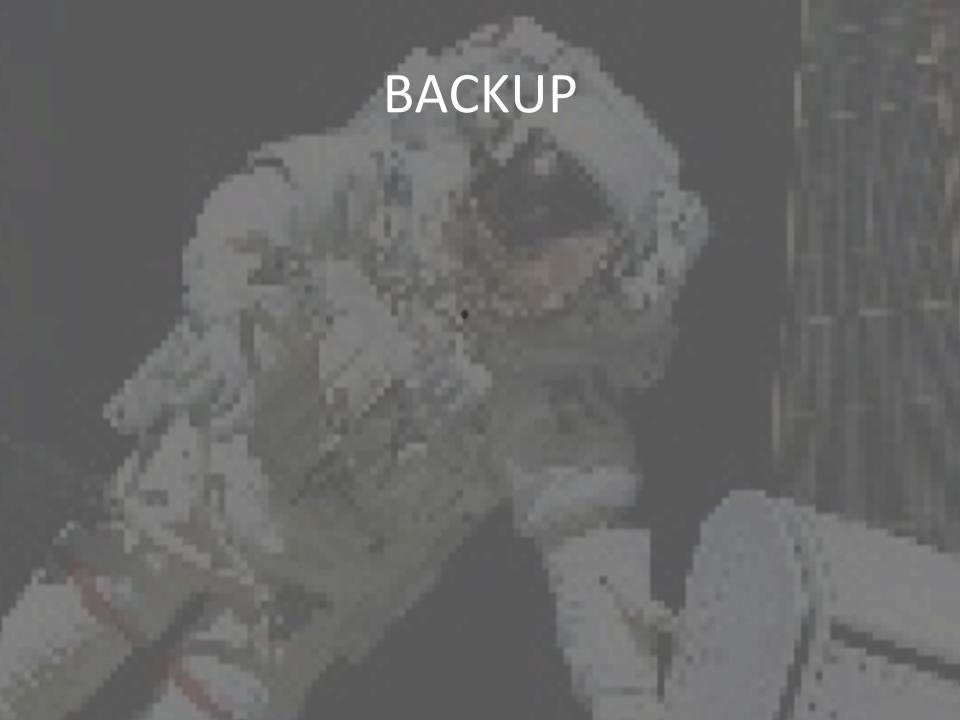




### Implications of the Worsening GCR Radiation Environment

- Direct observations (e.g., from LRO/CRaTER and MSL/Rad) critical!
- Deepest Solar Minimum and Weakest Maximum more than 80 years
  - Increased GCR radiation intensity, particularly in solar minima
  - Reduced allowable time in deep space for astronauts
  - Lower probability of SEP events
- Need Improved Understanding/Predictability of SEPs
  - Probability of Extreme Events
  - Resolve physics and predictability of extreme events

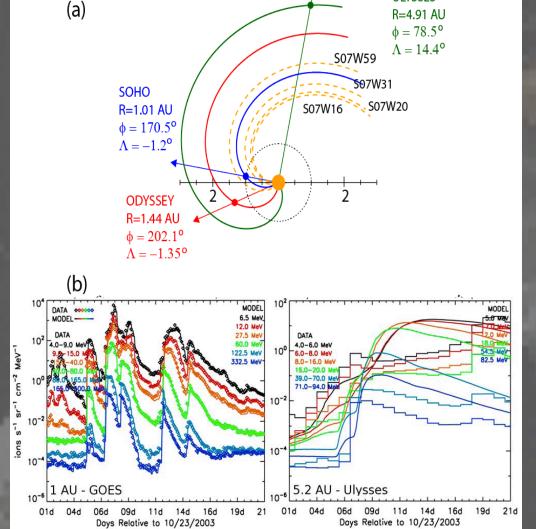


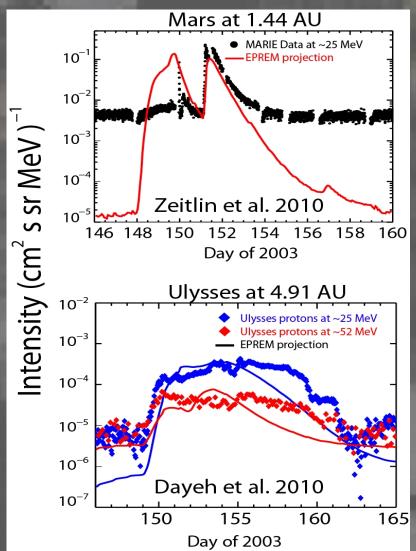


#### Connecting Events Through the Heliosphere

<u>Power of Distributed Observers:</u> Broad Longitudinal extent of SEP events has proved very useful for predicting SEP spectra and radiation doses at different locations in the inner heliosphere. Figures below show two recent papers by which SEP time profiles, onset, and radiation estimates were successfully predicted at Mars (Odyssey) and Ulysses located at 1.44 AU and 4.91 AU, respectively. 1 AU measurement from ACE, SoHO, and GOES

ULYSSES

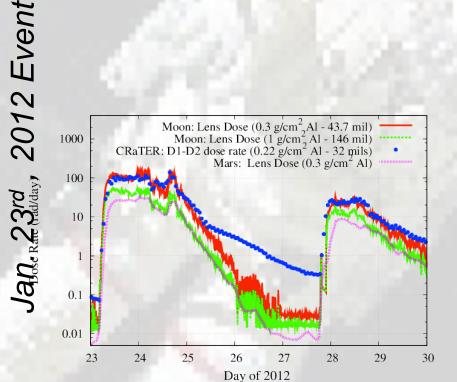


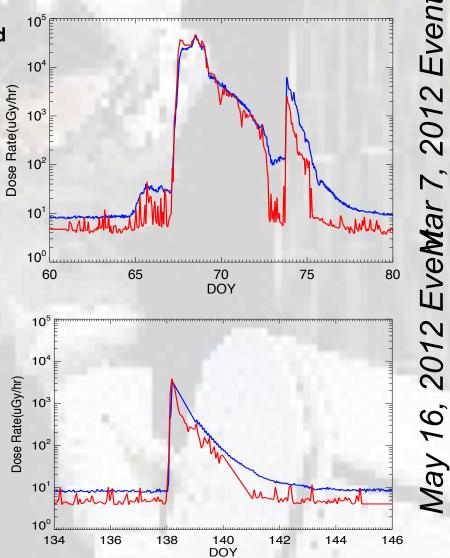


#### **SEP** events with PREDICCS

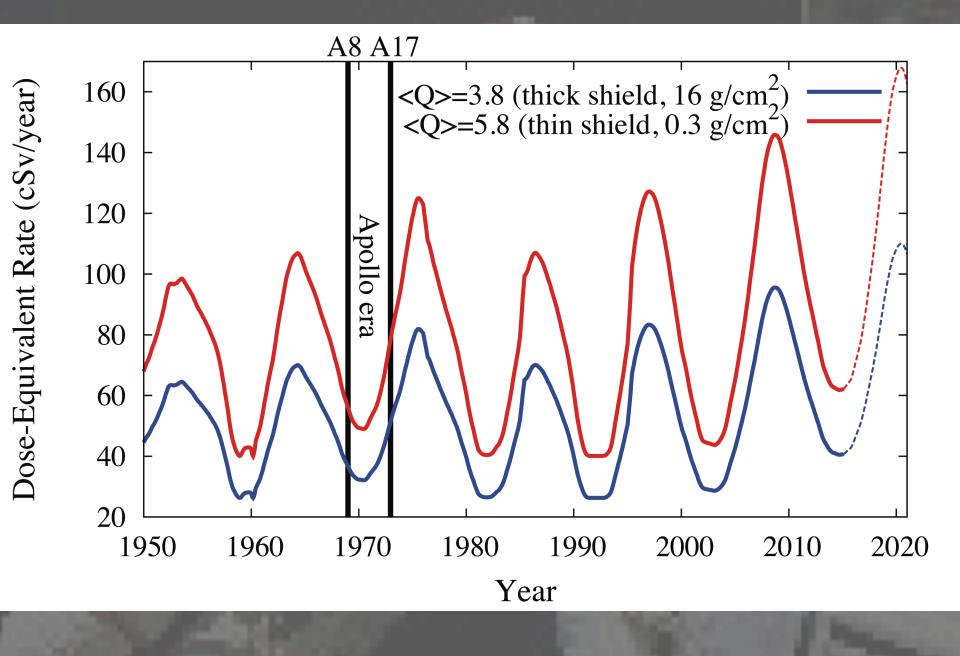
### **SEP Events During 2012: Indicators of Larger SEP Events in the New Cycle (24)**

- Shown here are the major SEP events of 2012 and the comparisons between CRaTER observations (blue) and prediccs predictions (red and green).
- Agreement reveals overall accuracy of models, while deviations likely reveal heavy ion contributions to dose observed by CRaTER

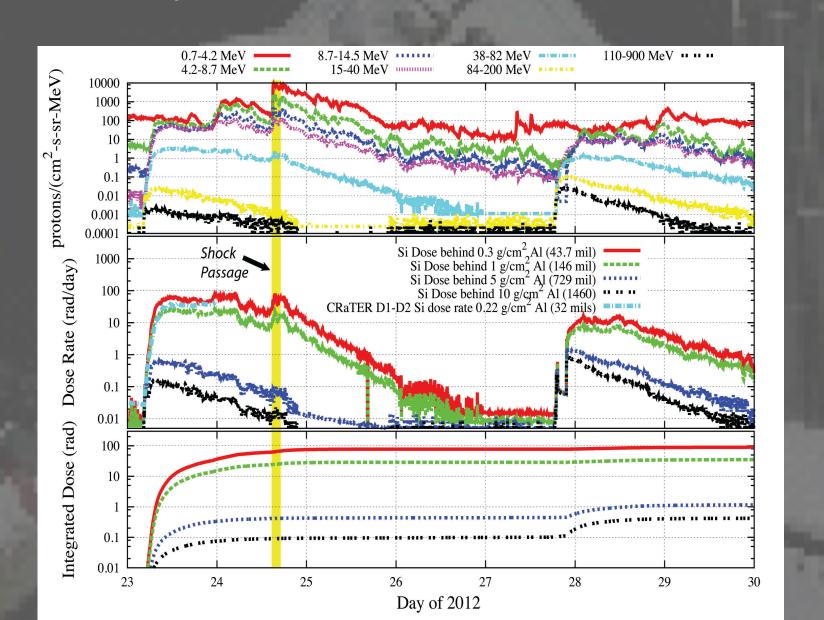


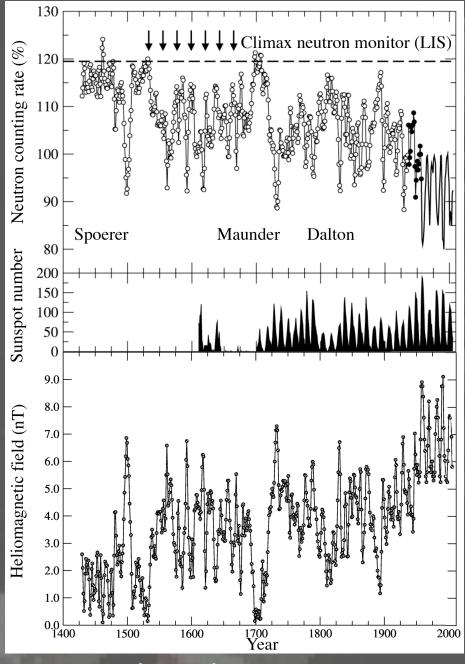


CRaTER (blue) EMMREM (red)



## Jan 23<sup>rd</sup> 2012 Event





- Galactic Cosmic Ray Record from Be10 in Ice Core (product of spallation)
- Do recent solar changes suggest larger changes in store?

### Halloween 2003 Storms

